

MATH 241

LECTURE 17

Higher Dimensional PDEs V (Spherical Harmonics)

In this lecture we focus on problems in spherical geometry, such as natural convection inside the earth (heat equation in a sphere) or seismic waves (wave eq.).

Of course, the real processes are more complex, but the main mathematical component of most models are these "simple" linear PDEs.

↳ A good understanding of separation of variables is very useful when one works with cylindrical or spherical harmonics (modified Bessel functions, Legendre...) in numerical methods.

↑ for physics or engineering people, most times you simply use the code from someone else.

A good understanding of what is expected to happen is necessary to check for error or to develop modifications/improvements.

- We are going to solve the wave equation with Dirichlet BCs, but basically, the same applies for other problems (heat eq., Laplace eq., ...).

• Problem: Wave Equation in a sphere with Dirichlet BC.

PDE: $u_{tt} = c^2 \Delta u$ in Ω

BC: $u = 0$ on Ω

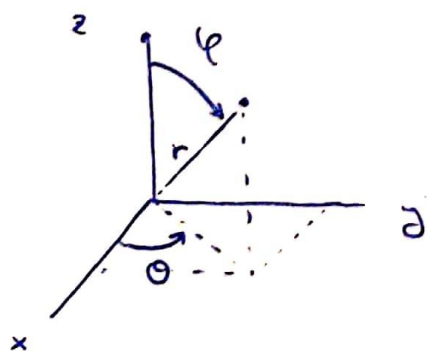
I.C: $u|_{t=0} = \alpha, u_t|_{t=0} = \beta$

} Ω sphere centered at origin with radius R .

Solution:

First, let's state the notation for the spherical coordinates:

[the notation changes from books and websites, so always be careful].



$$\left. \begin{aligned} x &= r \sin \varphi \cos \theta \\ y &= r \sin \varphi \sin \theta \\ z &= r \cos \varphi \end{aligned} \right\} \begin{aligned} -\pi &\leq \theta \leq \pi \\ 0 &\leq \varphi \leq \pi \\ 0 &\leq r \leq R \end{aligned}$$

↪ math typical notation.

Remark: In physics they usually use φ, θ in the contrary way.

In many books the radial variable is denoted ρ .

In the book (R. Haberman), ϕ is used instead of ψ (I don't like it: it's less common and we have been using ϕ a lot to denote a 2d function of spatial variables).

- Remember that we can separate the time and spatial variables without knowing the geometry of Ω :

$$u(r, \vartheta, \varphi, t) = \phi(r, \vartheta, \varphi) h(t) \leadsto$$

$$\leadsto \phi h'' = c^2 h \Delta \phi \rightarrow \frac{\Delta \phi}{\phi} = \frac{1}{c^2} \frac{h''}{h} = -\lambda \leadsto$$

$$\left. \begin{array}{l} \textcircled{1} \Delta \phi + \lambda \phi = 0 \text{ in } \Omega \\ \phi = 0 \text{ on } \partial \Omega \end{array} \right\} \quad \textcircled{2} h''(t) = -\lambda c^2 h(t)$$

where we have used that

$$u(r, \vartheta, \varphi, t) = 0 = \phi(r, \vartheta, \varphi) h(t) \text{ for all } t \geq 0 \Rightarrow \phi \equiv 0 \text{ on } \partial \Omega, \\ \text{and } (r, \vartheta, \varphi) \in \partial \Omega$$

Remark: In $\textcircled{1}$, Δ denotes the 3-d Laplacian. The formulas for integration by parts are the same in

2d and 3d, so we can easily prove that $\lambda > 0$.

Proof We do it here once:

$$\Delta\phi + \lambda\phi = 0 \rightarrow \int_{\Omega} \phi \Delta\phi \, dV = -\lambda \int_{\Omega} \phi^2 \, dV \rightarrow$$

$$\Rightarrow \lambda = \frac{-1}{\int_{\Omega} \phi^2 \, dV} \left(- \int_{\Omega} |\nabla\phi|^2 \, dV + \int_{\partial\Omega} \phi \, \vec{n} \cdot \nabla\phi \, dS \right) \geq 0.$$

BC.

If $\lambda = 0$, then $|\nabla\phi| = 0$ on $\Omega \Rightarrow \phi = c \rightarrow \phi = 0$ ~~*~~
 $\uparrow$
 $\phi = 0$ on $\partial\Omega$

$\rightarrow$ In summary, the Helmholtz equation with Dirichlet BC. has strictly positive eigenvalues, both in 2d and 3d ||

• Because $\lambda > 0$ (although unknown values yet),

$$h(t) = c_1 \cos(c\sqrt{\lambda} t) + c_2 \sin(c\sqrt{\lambda} t).$$

The values of λ will be given by solving the eigenvalue problem ①.

• Remark: 1) Problem ① will appear when solving the heat equation too.

2) For $\lambda = 0$, $\phi \equiv 0$ is the only solution.

3) If we had insulation BC, then $\lambda \geq 0$ and for $\lambda = 0$ we obtain $\phi \equiv c$.

4) If we have different BCs, the case $\lambda = 0$ might be interesting. It corresponds to solving the Laplace equation in a sphere (which gives the equilibrium sol. to the heat eq.).

Solving ① Helmholtz Equation in a sphere (Dirichlet BCs).

$$\begin{aligned} \Delta \phi + \lambda \phi &= 0 \quad \text{in } \Omega \\ \phi &= 0 \quad \text{on } \partial \Omega \end{aligned} \quad \left\{ \begin{array}{l} \Omega \equiv \text{sphere.} \end{array} \right.$$

In spherical coordinates,

$$\Delta \phi(r, \theta, \varphi) = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \phi}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial \phi}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 \phi}{\partial \varphi^2}$$

• Now we try to separate variables:

$\phi(r, \theta, \varphi) = F(r)G(\theta)H(\varphi)$ and substitute in the PDE,

$$\Delta \phi(r, \vartheta, \varphi)$$

$$\frac{G(\vartheta)H(\varphi)}{r^2} \frac{d}{dr} (r^2 F'(r)) + \frac{F(r)G(\vartheta)}{r^2 \sin \varphi} \frac{d}{d\varphi} (\sin \varphi H'(\varphi)) + \frac{F(r)H(\varphi)}{r^2 \sin^2 \varphi} G''(\vartheta) +$$

$$+ \lambda F(r)G(\vartheta)H(\varphi) = 0 \rightarrow \left\{ \begin{array}{l} \text{multiply by } r^2 \sin^2 \varphi, \\ \text{divide by } FHG \end{array} \right\}$$

$$\frac{\sin^2 \varphi}{F(r)} \frac{d}{dr} (r^2 F'(r)) + \frac{\sin \varphi}{H(\varphi)} \frac{d}{d\varphi} (\sin \varphi H'(\varphi)) + \frac{G''(\vartheta)}{G(\vartheta)} + \lambda r^2 \sin^2 \varphi = 0 \rightarrow$$

$$\Rightarrow \left. \begin{array}{l} \textcircled{1.1} \quad G''(\vartheta) = -\eta G(\vartheta) \\ \text{+ periodic BC.} \end{array} \right\} \quad (\vartheta \equiv \text{horizontal angle, } -\bar{\alpha} \leq \vartheta \leq \bar{\alpha})$$

$$\textcircled{1.2} \quad \frac{1}{F(r)} \frac{d}{dr} (r^2 F'(r)) + \frac{1}{\sin \varphi H(\varphi)} \frac{d}{d\varphi} (\sin \varphi H'(\varphi)) + \lambda r^2 = -\frac{\eta}{\sin^2 \varphi}$$

• The BC $\phi = 0$ on ∂R translates to:

$$\phi(R, \vartheta, \varphi) = F(R)G(\vartheta)H(\varphi) = 0 \text{ for all } \vartheta, \varphi \Rightarrow F(R) = 0.$$

Remark: ① is a 3d problem, so we will need two separation constants (not counting λ).

• The solution to (1.1) is well-known:

$$\eta_n = n^2, n \geq 0, \text{ with } G(\theta) = c_1 \cos(n\theta) + c_2 \sin(n\theta).$$

• With $\eta_n = n^2$, we proceed to solve (1.2). We rewrite it trying to separate r and φ :

$$(1.2) \quad \frac{1}{F(r)} \frac{d}{dr} (r^2 F'(r)) + \lambda r^2 = - \frac{1}{\sin \varphi H(\varphi)} \frac{d}{d\varphi} (\sin \varphi H'(\varphi)) + \frac{n^2}{\sin^2 \varphi} = \mu$$

which gives rise to:

$$(1.2.1) \quad \left. \begin{aligned} \frac{d}{d\varphi} (\sin \varphi H'(\varphi)) + \left(\mu \sin \varphi - \frac{n^2}{\sin \varphi} \right) H(\varphi) &= 0 \end{aligned} \right\}$$

$$(1.2.2) \quad \left. \begin{aligned} \frac{d}{dr} (r^2 F'(r)) + (\lambda r^2 - \mu) F(r) &= 0 \\ F(R) &= 0 \end{aligned} \right\}$$

• We choose to solve first (1.2.1) because it doesn't depend on λ . This problem will determine the values of μ . Then, with those values of μ , we will use (1.2.2) to find the values of λ .

→ It is clear then for (1.2.1) and (1.2.2) there must be some "hidden" boundary conditions.

- Solving (1.2.1) : Legendre Polynomials and Associated Legendre functions.

The linear second-order ODE (1.2.1) is of Sturm-Liouville type with $p(\varphi) = \sin \varphi$, $\sigma(\varphi) = \sin \varphi$, $q(\varphi) = -\frac{n^2}{\sin \varphi}$.

It is not a regular one due to the points $\varphi = 0$ and $\varphi = \pi$, which corresponds to the North and South poles.

Notice that this is simply a singularity coming from the spherical coordinates (like $r=0$ in plane coord.).

To have meaningful solutions, we must ensure that the solutions are bounded at those points:

$$|H(0)| < \infty, |H(\pi)| < \infty$$

So (1.2.1) is the BVP given by

$$\begin{cases} \frac{d}{d\varphi}(\sin \varphi H'(\varphi)) + \left(\mu \sin \varphi - \frac{n^2}{\sin \varphi} \right) H(\varphi) = 0, & 0 < \varphi < \pi, \\ |H(0)| < \infty, |H(\pi)| < \infty \end{cases}$$

This is a new ODE, which in most book appears in a different form:

Change of variables: $x = \cos \theta$ (bijection from $[0, \pi]$ to $[-1, 1]$).

The equation becomes

$$\left\| \frac{d}{dx} \left((1-x^2) H'(x) \right) + \left(\mu - \frac{n^2}{1-x^2} \right) H(x) = 0 \right\| \text{Associated Legendre ODE.}$$

$-1 < x < 1$ [L]

Claim: The general solution to [L] is given by

$$H(x) = c_1 P_\alpha^n(x) + c_2 Q_\alpha^n, \text{ where } \mu = \alpha(\alpha+1) \text{ and}$$

P_α^n, Q_α^n are the associated Legendre functions of the first and second kind, respec.

More commonly known as Spherical Harmonics.

→ Both P_α^n, Q_α^n are generally unbounded at $x=1$ or $x=-1$.

The only ones that are bounded both at $x=1$ and $x=-1$ are the spherical harmonics of first kind with $\alpha = m$ an integer: (Hard to prove).

$$|P_m^n(1)| < \infty, |P_m^n(-1)| < \infty.$$

- We focus on the case $\alpha = m$ since we are only interested in bounded solutions.

It is surprising that the spherical harmonics of first kind with integer $\alpha = m$ are quite simple functions: they are polynomials!

1) Legendre polynomials: $P_m^0(x)$

These satisfy the ODE $\frac{d}{dx}((1-x^2)H'(x)) + m(m+1)H(x) = 0$ with $|H(1)| < \infty, |H(-1)| < \infty$

They satisfy Rodrigues's formula:

$$P_m^0(x) = \frac{1}{2^m n!} \frac{d^m}{dx^m} (x^2-1)^m \Rightarrow P_m^0 \text{ is a polynomial of degree } m.$$

2) Associated Legendre polynomials: $P_m^n(x)$

$$P_m^n(x) = (x^2-1)^{n/2} \frac{d^n P_m^0}{dx^n}$$

• Remark: Since P_m^0 is a polynomial of degree m ,

$$P_m^n(x) \equiv 0 \text{ for } n > m.$$

- Summary: The solution to the eigenvalue problem

$$\begin{cases} \frac{d}{dx}((1-x^2)H'(x)) + \left(\mu - \frac{n^2}{1-x^2}\right)H(x) = 0, & -1 < x < 1, \\ |H(-1)| < \infty, |H(1)| < \infty. \end{cases}$$

is given by

$$\begin{cases} \mu_{n,m} = m(m+1) \text{ with } m \geq n \\ H_{n,m}(x) = P_m^n(x) \end{cases}$$

- Back to (1.2.1) (page -201-): $H_{n,m}(\varphi) = P_m^n(\cos \varphi)$
 $[x = \cos \varphi]$

That is, the solution to the eigenvalue problem (1.2.1) is

$$\begin{cases} \mu_{n,m} = m(m+1) \text{ with } m \geq n, \\ H_{n,m}(\varphi) = P_m^n(\cos \varphi) \end{cases}$$

Remark: The orthogonality of these eigenfunctions is easy to prove from the Associated Legendre eq. (good exercise).

• Solving 1.2.2: Spherical Bessel Functions.

Now we know that $\mu = m(m+1)$, so we have to solve

$$\left. \begin{aligned} \frac{d}{dr} (r^2 F'(r)) + (\lambda r^2 - m(m+1)) F(r) &= 0, \quad 0 < r < R \\ F(R) &= 0, \quad |F(0)| < \infty \end{aligned} \right\}$$

↑ added as usual

This ODE reminds Bessel's ODE, so let's first do the same change to get rid of the λ dependence:

$$x = \sqrt{\lambda} r \quad (\text{we know that } \lambda > 0, \text{ so this is a bijection}).$$

$$\rightarrow \boxed{x^2 F''(x) + 2x F'(x) + (x^2 - m(m+1)) F(x) = 0} \quad \boxed{\text{Spherical Bessel ODE.}}$$

The general solution is

$$F(x) = c_1 j_m(x) + c_2 y_m(x), \text{ where}$$

$$j_m(x) = \frac{1}{\sqrt{x}} J_{m+\frac{1}{2}}(x)$$

$$y_m(x) = \frac{1}{\sqrt{x}} Y_{m+\frac{1}{2}}(x)$$

are called Spherical Bessel functions
(of the first and second kind).

(*) Bessel functions of order $m+\frac{1}{2}$ look like the ones we have studied before.

It is just a matter of substitution to check that j_m, y_m solve the Spherical Bessel ODE, using that $J_{m+\frac{1}{2}}, Y_{m+\frac{1}{2}}$ solve Bessel ODE.

That is, if $J_{m+\frac{1}{2}}, Y_{m+\frac{1}{2}}$ are solutions of $x^2 F''(x) + x F'(x) + (x^2 - (m+\frac{1}{2})^2) F(x) = 0$,

then j_m, y_m solve $x^2 F''(x) + 2x F'(x) + (x^2 - m(m+1)) F(x) = 0$ ||

Remark: The functions j_m, y_m behave and look very similar to J_m, Y_m ("same" graphs).

• Going back to r :

$$F(r) = c_1 j_m(\sqrt{\lambda} r) + c_2 y_m(\sqrt{\lambda} r)$$

$$|F(0)| < \infty \Rightarrow c_2 = 0$$

$$F(R) = 0 \Leftrightarrow j_m(\sqrt{\lambda} R) = 0 \Leftrightarrow J_{m+\frac{1}{2}}(\sqrt{\lambda} R) = 0 \Leftrightarrow$$

$$\Leftrightarrow \sqrt{\lambda_{m,\ell}} R = z_{m,\ell}, \text{ with } z_{m,\ell} \text{ the } \ell^{\text{th}} \text{ root of } J_{m+\frac{1}{2}}.$$

• Remark: (not needed) A nice formula for the spherical Bessel functions is

$$j_m(x) = \frac{1}{x} J_{m+\frac{1}{2}}(x) = x^m \left(-\frac{1}{x} \frac{d}{dx} \right)^m \left(\frac{\sin x}{x} \right).$$

In particular, $j_0(x)$ is the Sinc(x) function: $\frac{\sin x}{x}$. ⌋

• Joining all together:

(1.1) $\eta_n = n^2$, $G(\vartheta) = c_1 \cos(n\vartheta) + c_2 \sin(n\vartheta)$,

(1.2.1) $\mu_{n,m} = m(m+1)$ with $m \geq n$,

$$H_{n,m}(\varphi) = P_m^n(\cos \varphi).$$

(1.2.2) $\lambda_{m,e} = \left(\frac{z_{m,e}}{R} \right)^2$, with $z_{m,e}$ the e^{th} root of $J_{m+\frac{1}{2}}$,

$$F(r) = \frac{1}{r} J_{m+\frac{1}{2}}(r) \equiv j_m(r).$$

Superposition:

$$u(r, \vartheta, \varphi, t) = \sum_{n=0}^{\infty} \sum_{m=n}^{\infty} \sum_{e=1}^{\infty} \left(A_{n,m,e} \cos(c \sqrt{\lambda_{m,e}} t) + B_{n,m,e} \sin(c \sqrt{\lambda_{m,e}} t) \right).$$

$$\cdot \left(C_{n,m,e} \cos(n\vartheta) + D_{n,m,e} \sin(n\vartheta) \right) j_m(r) P_m^n(\cos \varphi).$$

I.C.:

$$u(r, \vartheta, \varphi, 0) = \alpha(r, \vartheta, \varphi)$$

$$u_t(r, \vartheta, \varphi, 0) = \beta(r, \vartheta, \varphi) \quad \leadsto A, B, C, D //$$